Quantitative Finance

Lectures in Quantitative Finance Spring Term 2022

1. Basic probability theory & regression



Prof. Dr. Erich Walter Farkas walter.farkas@bf.uzh.ch





Basic probability theory & regression





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1. Basic probability theory

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Basic probability theory

The very base of our quantitative studies is a *probability space* $(\Omega, \mathcal{F}, \mathbb{P})$ consisting of

- ightharpoonup a sample space $\Omega \neq \emptyset$,
- ightharpoonup a σ -algebra $\mathcal{F}\subset\mathcal{P}(\Omega)$, a subset of the power set of Ω , and

The sample space Ω is an arbitrary non-empty set. The only assumption here is *non-emptiness*, but it is important and needs to be stated!

The elements of Ω are usually denoted by ω for the simple reason that 'small omegas' live in the 'big omega'. We may give them subscripts, bars, hats, and so on, $\omega_1, \omega_k, \bar{\omega}, \hat{\omega}$.

In principle, you are free to use your imagination at the cost of transparency: $\odot,\star,\odot\in\Omega$. However, we strongly advise you against doing that. Some conventions are so firmly established that any deviation causes confusion.

By $\mathcal{P}(\Omega)$ or 2^{Ω} we denote the *power set* of Ω . This is the set of all subsets of Ω , including the empty set \emptyset and Ω .

Example. Consider the set $S = \{a, b, c\}$. All possible strict subsets are $\{a\}, \{b\}, \{c\}, \{a, b\}, \{a, c\}, \{b, c\}$. Including the empty set and S itself, we get the power set of S,

$$\mathcal{P}(S) = \{\emptyset, \{a\}, \{b\}, \{c\}, \{a, b\}, \{a, c\}, \{b, c\}, \{a, b, c\}\}\}.$$

Notice that S has |S|=3 elements and $|\mathcal{P}(S)|=8=2^3$. This is true in general for any *finite set*: if $|S|=n\in\mathbb{N}$, then $|\mathcal{P}(S)|=2^n$ (exercise). That is the reasoning behind the notation $\mathcal{P}(S)=2^S$.

Let Ω be some non-empty set. A σ -algebra or σ -field \mathcal{F} over Ω is a subset of the power set $\mathcal{P}(\Omega)$ which satisfies the following three properties:

- $\triangleright \Omega \in \mathcal{F}$.
- ightharpoonup if $A \in \mathcal{F}$, then $A^{\complement} := \Omega \setminus A := \{\omega \in \Omega \mid \omega \notin A\} \in \mathcal{F}$ (we say \mathcal{F} is closed under complementation),
- ightharpoonup if $A_1,A_2,\ldots\in\mathcal{F}$, then $\cup_{k=1}^{\infty}A_k\in\mathcal{F}$ (we say \mathcal{F} is closed under countable unions).

Exercise. Using the defining properties above, show that

- $\triangleright \emptyset \in \mathcal{F}$. and
- \triangleright if $A_1, A_2, \ldots \in \mathcal{F}$, then $\bigcap_{k=1}^{\infty} A_k \in \mathcal{F}$ (use De Morgan's laws).

The σ -algebra is just a definition of which sets may be considered as events.

Elements not in \mathcal{F} simply have no defined probability measure.

Why are they important?

- \triangleright It is not always possible to assign a measure to **all** subsets of Ω ,
- \triangleright Keeping track of σ -algebras allows us to formulate some key concepts in probability very elegantly!

Let Ω be a non-empty set and let \mathcal{F} be a σ -algebra over Ω .

We call a function $\mu: \mathcal{F} \to \overline{\mathbb{R}} = \mathbb{R} \cup \{-\infty, +\infty\}$ a *measure* if it satisfies

- $\triangleright \mu(\emptyset) = 0$,
- ightharpoonup non-negativity: for all $F \in \mathcal{F}$ we have $\mu(F) \geq 0$, and
- ho countable additivity (σ -additivity): for any countable collection $\{F_k\}_{k=1}^{\infty} \subset \mathcal{F}$ of pairwise disjoint sets we have $\mu(\bigcup_{k=1}^{\infty} F_k) = \sum_{k=1}^{\infty} \mu(F_k)$.

In general, we call a set together with a σ -algebra over that set, in our case the pair (Ω, \mathcal{F}) , a *measurable space*, since we can define a measure on it.

Elements of \mathcal{F} are called *measurable sets*.

The triplet $(\Omega, \mathcal{F}, \mu)$ is called a *measure space*.

Let Ω be a non-empty set and let \mathcal{F} be a σ -algebra over Ω .

We call a function $\mathbb{P}:\mathcal{F}\to[0,1]$ a probability measure if it satisfies

- $hd \mathbb{P}[\emptyset] = 0$ and $\mathbb{P}[\Omega] = 1$, and
- \triangleright \mathbb{P} is countably additive.

Notice that a probability measure is a special kind of measure with total measure of one.

For a probability measure \mathbb{P} , we call the triplet $(\Omega, \mathcal{F}, \mathbb{P})$ a *probability* space.

Let $(\Omega, \mathcal{F}, \mathbb{P})$ be a probability space and let $\mathcal{B}(\mathbb{R})$ be the *Borel* σ -algebra on \mathbb{R} . This is (without going into much detail here) the smallest σ -algebra on \mathbb{R} which *contains all the intervals*.

In particular, sets of the form $(-\infty, x]$ for some $x \in \mathbb{R}$ are elements of $\mathcal{B}(\mathbb{R})$.

A real-valued random variable is a measurable function $X : \Omega \to \mathbb{R}$. This means that for every set $B \in \mathcal{B}(\mathbb{R})$ the preimage $X^{-1}(B) := \{\omega \in \Omega \mid X(\omega) \in B\}$ is an element of \mathcal{F} .

To emphasise the dependence of measurability on the underlying σ -algebra, authors also write $X:(\Omega,\mathcal{F})\to(\mathbb{R},\mathcal{B}(\mathbb{R})).$

The *probability* that X takes on a value in a specified measurable set $B \in \mathcal{B}(\mathbb{R})$ is written as

$$\mathbb{P}[\{\omega \in \Omega \mid X(\omega) \in B\}].$$

Most of the time, we will abbreviated the set $\{\omega \in \Omega \mid X(\omega) \in B\}$ by $\{X \in B\}$ and hence, we will write $\mathbb{P}[X \in B]$.

Notice that $X^{-1}(B) = \{X \in B\}$ and that already here the notion of measurability is essential: the set $\{X \in B\}$ has to be an element of \mathcal{F} so that we can measure it!

Every *real-valued* random variable X can be described by its *cumulative* distribution function (cdf) $F_X \colon \mathbb{R} \to [0,1]$ defined as

$$F_X(x) = \mathbb{P}[X \leq x]$$
.

Notice that, as mentioned above, $\{X \leq x\} = X^{-1}((-\infty, x]) \in \mathcal{F}$.

Exercise. Argue that any cdf *F* is *increasing* and that

$$\lim_{x \to -\infty} F(x) = 0$$
 and $\lim_{x \to +\infty} F(x) = 1$.

We call a real-valued random variable $X : \Omega \to S$ for some $S \subset \mathbb{R}$ discrete if its image is *finite* or *countably infinite*.

The function $f_X \colon S \to [0,1]$ which gives the probability that X is exactly equal to a certain value in its image is called *probability mass function* (pmf) (not pdf!),

$$f_X(x) = \mathbb{P}[X = x]$$
 for $x \in S$.

Notice that we can assume without loss of generality (w.l.o.g.) that these probabilities are non-zero if we restrict S to the image of X.

Since the image of X is at most countably infinite, we can also enumerate the corresponding probabilities. That means for every x_k in the image of X we can define $p_k = \mathbb{P}[X = x_k] = f_X(x_k)$.

Exercise. Argue that $\sum_{x \in S} f_X(x) = \sum_{k=1}^{\infty} p_k = 1$.

Exercise. Let $\Omega = \{\omega_1, \dots, \omega_6\}$ describe all possible states after rolling a six faced dice, that is $\omega_1 = \{\text{After rolling the dice, it shows '1'}\}, \dots, \omega_6 = \{\text{After rolling the dice, it shows '6'}\}.$

The discrete random variable which gives us the *value of the dice* is defined as $X(\omega_1) = 1, \dots, X(\omega_6) = 6$.

We assume that the dice is perfect, that means every face has the same probability, $\mathbb{P}[\omega_1] = \ldots = \mathbb{P}[\omega_6] = \frac{1}{6}$.

Write down the cumulative distribution function $F_X(x) = \mathbb{P}[X \le x]$ of X. Recall that $\{X \le x\} = \{\omega \in \Omega \mid X(\omega) \le x\}$. We say X is a *continuous random variable* if its cdf is continuous. In particular, this implies that the image of X is *uncountably infinite*.

If X is absolutely continuous, then X admits a probability density function (pdf) f_X which assigns probabilities to intervals.

We have

$$\mathbb{P}[a \le X \le b] = \int_a^b f_X(x) dx,$$

in particular, we have

$$F_X(b) = \int_{-\infty}^b f_X(x) \mathrm{d}x.$$

Example. Let X be (continuous) uniformly distributed on the closed interval [a,b] for $-\infty < a < b < \infty$. We write $X \sim \mathcal{U}(a,b)$. By definition, the pdf is given by

$$f_X(x) = \begin{cases} \frac{1}{b-a} & \text{for } a \le x \le b \,, \\ 0 & \text{else.} \end{cases}$$

Hence, the cdf is given as

$$F_X(x) = \int_{-\infty}^x f_X(t) dt = \begin{cases} 0 & \text{for } x < a, \\ \frac{x-a}{b-a} & \text{for } a \le x \le b, \\ 1 & \text{for } x > b. \end{cases}$$

Exercise. Let X be a absolutely continuous random variable with cdf F_X and pdf f_X .

- ightharpoonup Show that $F_X'(x) = f_X(x)$ (use the first fundamental theorem of calculus).
- ▷ Argue that $\mathbb{P}[X = x] = 0$ for any $x \in \mathbb{R}$.

The expected value (or expectation, mean, average, first moment) of a random variable X intuitively characterises the central tendency or long-run average value.

For absolutely continuous random variables it is defined via the integral

$$\mathbb{E}[X] = \int_{-\infty}^{\infty} x \, f_X(x) \, \mathrm{d}x.$$

In the discrete case, the integral becomes a sum

$$\mathbb{E}[X] = \sum_{k=1}^{\infty} x_k \mathbb{P}[X = x_k] = \sum_{k=1}^{\infty} x_k p_k.$$

Example. Let $X \sim \mathcal{U}(a, b)$, then the expected value of X is

$$\mathbb{E}[X] = \int_{-\infty}^{\infty} x \, f_X(x) \, \mathrm{d}x = \int_a^b \frac{x}{b-a} \mathrm{d}x = \frac{x^2}{2(b-a)} \Big|_{x=a}^b$$
$$= \frac{b^2 - a^2}{2(b-a)} = \frac{a+b}{2} \, .$$

Now let X denote the value of a perfect dice as in the example above. The expected value of X is

$$\mathbb{E}[X] = \sum_{k=1}^{\infty} x_k \mathbb{P}[X = x_k] = \sum_{k=1}^{6} k \frac{1}{6} = \frac{21}{6} = 3.5.$$

Exercise. Show that the expectation is a *linear functional*. This means that for any random variables X, Y and any real number $c \in \mathbb{R}$ we have

$$riangleright \mathbb{E}[X+Y] = \mathbb{E}[X] + \mathbb{E}[Y]$$
, and

$$\triangleright \mathbb{E}[cX] = c\mathbb{E}[X].$$

The *variance* of a random variable is the expectation of the squared deviation from its mean. It is defined as

$$\operatorname{Var}(X) = \mathbb{V}[X] = \mathbb{E}[(X - \mathbb{E}[X])^2].$$

The standard deviation (std) $\sigma(X)$ of a random variable X is the square root of its variance,

$$\sigma(X) = \sqrt{\mathbb{V}[X]}.$$

Exercise. Let X be a random variable and let $c \in \mathbb{R}$. Show that

- ho $\mathbb{V}[X] = \mathbb{E}[X^2] \mathbb{E}[X]^2$ (use linearity of the expected value),
- $\triangleright \mathbb{V}[X+c] = \mathbb{V}[X],$
- $\triangleright \ \mathbb{V}[cX] = c^2 \mathbb{V}[X].$

Find an example of two random variables X, Y such that

$$\mathbb{V}[X + Y] \neq \mathbb{V}[X] + \mathbb{V}[Y]$$
.

Hint: Make it simple by taking for example $\Omega = \{\omega_1, \omega_2\}$ and defining $X(\omega_1) = 1$ and $X(\omega_2) = 0$. What could be a possible definition for Y?

Example. Let $X \sim \mathcal{U}(a, b)$, then the variance of X is

$$\mathbb{V}[X] = \mathbb{E}[X^2] - \mathbb{E}[X]^2 = \int_a^b \frac{x^2}{b-a} dx - \left(\frac{a+b}{2}\right)^2$$
$$= \frac{x^3}{3(b-a)} \Big|_{x=a}^b - \left(\frac{a+b}{2}\right)^2 = \frac{b^3 - a^3}{3(b-a)} - \left(\frac{a+b}{2}\right)^2.$$

We can now expand the fraction so that we have the same denominator and get

$$\mathbb{V}[X] = \frac{4(b^3 - a^3) - 3(b - a)(a + b)^2}{12(b - a)}$$
$$= \frac{b^3 + 3ba^2 - 3ab^2 - a^3}{12(b - a)} = \frac{(b - a)^3}{12(b - a)} = \frac{(b - a)^2}{12}.$$

Example. Let X denote the value of a dice again. The variance of X is given by

$$V[X] = \mathbb{E}[X^2] - \mathbb{E}[X]^2 = \sum_{k=1}^6 k^2 \frac{1}{6} - \left(\frac{21}{6}\right)^2 = \frac{91}{6} - \left(\frac{21}{6}\right)^2$$
$$= \frac{35}{12} = 2.91\overline{6}.$$

Example. One of the most prominent continuous probability distributions is the *normal or Gauss distribution*. We say a random variable X is *normally distributed* with *mean (or location parameter)* μ and *variance (or scale parameter)* σ^2 and we write $X \sim \mathcal{N}(\mu, \sigma^2)$ if its probability density function is given by

$$f_X(x) = \frac{1}{\sqrt{2\pi\sigma^2}} \exp\left(-\frac{(x-\mu)^2}{2\sigma^2}\right)$$
.

We call an $X \sim \mathcal{N}(0,1)$ a *standard* normal random variable.

For the density function of a standard normal random variable, by convention, we often use the symbol φ ,

$$\varphi(x) = \frac{1}{\sqrt{2\pi}} \exp\left(-\frac{x^2}{2}\right) ,$$

and for its cumulative distribution function we use Φ .

Notice that

$$f_X(x) = \frac{1}{\sigma} \varphi\left(\frac{x-\mu}{\sigma}\right) .$$

Exercise. Let $X \sim \mathcal{N}(\mu, \sigma^2)$. We call the location parameter μ the mean and the scale parameter σ^2 the variance of X. But are they really what their names suggest? Let's find out.

First of all, you have to either accept the fact, that

$$\int_{-\infty}^{\infty} e^{\frac{-x^2}{2\sigma^2}} dx = \sqrt{2\pi\sigma^2},$$

or ask in the break or after the lecture how you can calculate this.

So we want to show that

$$\mathbb{E}[X] = \frac{1}{\sqrt{2\pi\sigma^2}} \int_{-\infty}^{\infty} x e^{-\frac{(x-\mu)^2}{2\sigma^2}} dx = \mu.$$

We start with the substitution $z=x-\mu$ with $\mathrm{d}z=\mathrm{d}x$. We notice that the limits stay the same.

Hence, we get

$$\mathbb{E}[X] = \frac{1}{\sqrt{2\pi\sigma^2}} \int_{-\infty}^{\infty} (z+\mu) e^{-\frac{z^2}{2\sigma^2}} dz$$

$$= \frac{1}{\sqrt{2\pi\sigma^2}} \left(\underbrace{\int_{-\infty}^{\infty} z e^{-\frac{z^2}{2\sigma^2}} dz + \mu \int_{-\infty}^{\infty} e^{-\frac{z^2}{2\sigma^2}} dz}_{=\sqrt{2\pi\sigma^2}} \right) = \mu.$$

Notice that the first integral evaluates to 0, since $ze^{-\frac{z^2}{2\sigma^2}}$ is an *odd* function.

$$\mathbb{V}[X] = \mathbb{E}[X^2] - \mathbb{E}[X]^2 = \mathbb{E}[X^2] - \mu^2 = \sigma^2.$$

So we have to calculate $\mathbb{E}[X^2]$. Using the same substitution as above, we get

$$\mathbb{E}[X^{2}] = \frac{1}{\sqrt{2\pi\sigma^{2}}} \int_{-\infty}^{\infty} (z+\mu)^{2} e^{-\frac{z^{2}}{2\sigma^{2}}} dz$$

$$= \frac{1}{\sqrt{2\pi\sigma^{2}}} \left(\mu^{2} \underbrace{\int_{-\infty}^{\infty} e^{-\frac{z^{2}}{2\sigma^{2}}} dz}_{=\sqrt{2\pi\sigma^{2}}} + 2\mu \underbrace{\int_{-\infty}^{\infty} z e^{-\frac{z^{2}}{2\sigma^{2}}} dz}_{=0} + \int_{-\infty}^{\infty} z^{2} e^{-\frac{z^{2}}{2\sigma^{2}}} dz \right)$$

$$= \mu^{2} + \frac{1}{\sqrt{2\pi\sigma^{2}}} \int_{-\infty}^{\infty} z^{2} e^{-\frac{z^{2}}{2\sigma^{2}}} dz.$$

Hence, we can already intuit that the second term must evaluate to σ^2 .

To show this, we want use a second substitution $x=z^2$ with $\mathrm{d} x=2z\mathrm{d} z$. However, since we are integrating from $-\infty$, we cannot directly do that. Instead, we notice that the integrand is an *even function* and thus, we can split the integral in two equal parts at 0. We get

$$\int_{-\infty}^{\infty} z^2 e^{-\frac{z^2}{2\sigma^2}} dz = 2 \int_{0}^{\infty} z^2 e^{-\frac{z^2}{2\sigma^2}} dz = 2 \int_{0}^{\infty} x e^{-\frac{x}{2\sigma^2}} \frac{1}{2\sqrt{x}} dx$$
$$= \int_{0}^{\infty} \sqrt{x} e^{-\frac{x}{2\sigma^2}} dx.$$

Next, we use integration by parts to get

$$\int_0^\infty \sqrt{x} e^{-\frac{x}{2\sigma^2}} dx = -2\sigma^2 e^{-\frac{x}{2\sigma^2}} \sqrt{x} \Big|_{x=0}^\infty + \sigma^2 \int_0^\infty \frac{1}{\sqrt{x}} e^{-\frac{x}{2\sigma^2}} dx.$$

Notice that the first term is equal to 0.

Finally, we substitute again with $x = z^2$, dx = 2zdz and arrive at

$$\sigma^2 \int_0^\infty \frac{1}{\sqrt{x}} e^{-\frac{x}{2\sigma^2}} \mathrm{d}x = \sigma^2 \int_0^\infty \frac{1}{z} 2z e^{-\frac{z^2}{2\sigma^2}} \mathrm{d}z = \sigma^2 \sqrt{2\pi\sigma^2} \,.$$

Plugging everything in, we get

$$V[X] = \mathbb{E}[X^2] - \mathbb{E}[X]^2 = \mu^2 + \sigma^2 - \mu^2 = \sigma^2.$$

We conclude that calling the location parameter μ 'mean' and the scale parameter σ^2 'variance' is justified.

With the above results we can easily show that any normally distributed random variable $X \sim \mathcal{N}(\mu, \sigma^2)$ can be written in terms of a standard normally distributed random variable $Z \sim \mathcal{N}(0,1)$ as

$$X = \mu + \sigma Z$$
.

Indeed, we have

$$\mathbb{E}[\mu + \sigma Z] = \mu + \sigma \mathbb{E}[Z] = \mu = \mathbb{E}[X]$$

and

$$\mathbb{V}[\mu + \sigma Z] = \sigma^2 \mathbb{V}[Z] = \sigma^2 = \mathbb{V}[X].$$

The *covariance* between two real-valued random variables X and Y (with finite second moment) is the expected value of the product of their deviations from their expected values,

$$cov(X, Y) = \sigma_{XY} = \sigma(X, Y) = \mathbb{E}[(X - \mathbb{E}[X])(Y - \mathbb{E}[Y])].$$

Using the linearity of the expected value, we can simplify it to

$$cov(X, Y) = \mathbb{E}[(X - \mathbb{E}[X])(Y - \mathbb{E}[Y])]$$

$$= \mathbb{E}[XY - X\mathbb{E}[Y] - \mathbb{E}[X]Y + \mathbb{E}[X]\mathbb{E}[Y]]$$

$$= \mathbb{E}[XY] - \mathbb{E}[X]\mathbb{E}[Y].$$

Pearson's correlation coefficient is a measure of the linear correlation between two (non-constant) random variables X and Y, and it is defined as

$$\rho_{X,Y} = \frac{\operatorname{cov}(X,Y)}{\sqrt{\mathbb{V}[X]}\sqrt{\mathbb{V}[Y]}} = \frac{\sigma_{X,Y}}{\sigma_X\sigma_Y}.$$

If two random variables are independent, then they are also (linearly) uncorrelated. It is a common misconception that the reverse is also true. Uncorrelatedness does not imply independence! Bear this in mind, we will define independence below.

Exercise. Let X and Y be two random variables. Show that Pearson's correlation coefficient always lies in the interval [-1,1]. Use the Cauchy-Schwarz inequality.

We have already seen, that the expected value is linear and that the variance, in general, is not. Indeed, the variance of the sum of two random variables is given by

$$V[X + Y] = \mathbb{E}[(X + Y)^2] - \mathbb{E}[X + Y]^2$$

$$= \mathbb{E}[X^2 + 2XY + Y^2] - \mathbb{E}[X]^2 - 2\mathbb{E}[X]\mathbb{E}[Y] - \mathbb{E}[Y]^2$$

$$= \mathbb{E}[X^2] - \mathbb{E}[X^2] + \mathbb{E}[Y^2] - \mathbb{E}[Y]^2 + 2(\mathbb{E}[XY] - \mathbb{E}[X]\mathbb{E}[Y])$$

$$= V[X] + V[Y] + 2\text{cov}(X, Y).$$

Note that the variance of a sum is the sum of variances only if the variables are uncorrelated.

We can extend these results to a general weighted sum of random variables. Let $\{X_k\}_{k=1}^n$ be a family of random variables. Define

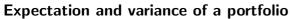
$$Y = \sum_{k=1}^{n} w_k X_k .$$

This could represent a portfolio of assets, where X_k represents, for example, the return of asset k. We assume that $\sum_{k=1}^{n} w_k = 1$, that means we invest all our wealth in the portfolio.

By linearity of the expected value, we immediately get that the expected value of the weighted sum is the weighted sum of the expected values,

$$\mathbb{E}[Y] = \mathbb{E}\Big[\sum_{k=1}^n w_k X_k\Big] = \sum_{k=1}^n \mathbb{E}[w_k X_k] = \sum_{k=1}^n w_k \mathbb{E}[X_k].$$

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Calculating the variance, we need to be more careful.

$$\mathbb{V}[Y] = \mathbb{V}\left[\sum_{k=1}^{n} w_{k} X_{k}\right] = \mathbb{E}\left[\left(\sum_{k=1}^{n} w_{k} X_{k}\right)^{2}\right] - \mathbb{E}\left[\sum_{k=1}^{n} w_{k} X_{k}\right]^{2}$$

$$= \mathbb{E}\left[\sum_{k=1}^{n} \sum_{\ell=1}^{n} w_{k} w_{\ell} X_{k} X_{\ell}\right] - \sum_{k=1}^{n} \sum_{\ell=1}^{n} w_{k} w_{\ell} \mathbb{E}[X_{k}] \mathbb{E}[X_{\ell}]$$

$$= \sum_{k=1}^{n} \sum_{\ell=1}^{n} w_{k} w_{\ell} \left(\mathbb{E}[X_{k} X_{\ell}] - \mathbb{E}[X_{k}] \mathbb{E}[X_{\ell}]\right).$$

$$= \sum_{k=1}^{n} \sum_{\ell=1}^{n} w_{k} w_{\ell} \operatorname{cov}(X_{k}, X_{\ell})$$

$$= \sum_{k=1}^{n} w_{k}^{2} \mathbb{V}[X_{k}] + 2 \sum_{k < \ell} w_{k} w_{\ell} \operatorname{cov}(X_{k}, X_{\ell}).$$

Since this is a lot to write, we want to abbreviate it and use matrix notation. For this we define the *weight vector* $w = (w_1, \dots, w_n)^T$ and the *covariance matrix*

$$\Sigma = \begin{pmatrix} \operatorname{cov}(X_1, X_1) & \operatorname{cov}(X_1, X_2) & \cdots & \operatorname{cov}(X_1, X_n) \\ \operatorname{cov}(X_2, X_1) & \operatorname{cov}(X_2, X_2) & \cdots & \operatorname{cov}(X_2, X_n) \\ \vdots & \vdots & \ddots & \vdots \\ \operatorname{cov}(X_n, X_1) & \operatorname{cov}(X_n, X_2) & \cdots & \operatorname{cov}(X_n, X_n) \end{pmatrix}$$

Now we can compactly write the variance of the sum as

$$\mathbb{V}[Y] = w^{\mathsf{T}} \Sigma w.$$

Exercise. Make sure that $w^{\mathsf{T}} \Sigma w$ is indeed equal to $\sum_{k=1}^{n} \sum_{\ell=1}^{n} w_k w_\ell \operatorname{cov}(X_k, X_\ell)$.

The *skewness* of a real-valued random variable is a measure of the asymmetry of the distribution about its mean. It is defined as the third standardised moment

$$\gamma_1(X) = \mathbb{E}\left[\left(\frac{X - \mathbb{E}[X]}{\sqrt{\mathbb{V}[X]}}\right)^3\right] = \mathbb{E}\left[\left(\frac{X - \mu}{\sigma}\right)^3\right].$$

We differentiate between

- negative skewness: long left tail (high probability of observing large negative values), and
- positive skewness: long right tail (high probability of observing large positive values).

The *kurtosis* of a real-valued random variable is (roughly speaking) a measure of the width of its tails. It is defined as the fourth standardised moment,

$$\operatorname{Kurt}[X] = \gamma_2(X) = \kappa(X) = \mathbb{E}\left[\left(\frac{X - \mu}{\sigma}\right)^4\right].$$

- A distribution with kurtosis of 3 is considered average and it is called mesokurtic. Any normally distributed random variable has kurtosis of 3. In fact, excess kurtosis is defined as kurtosis minus 3.
- A distribution with kurtosis greater than 3 is called *leptokurtic*. It indicates fatter tails.
- A distribution with kurtosis smaller than 3 is called *platykurtic*. It indicates thinner tails.

Exercise. As mentioned above, the kurtosis of any normally random variable is equal to 3. Using the fact that $\frac{X-\mu}{\sigma} \sim \mathcal{N}(0,1)$ and the 'tricks' we used to calculate the expectation and variance, prove this result.

Hint: First use the substitution $z = x^2$, then use integration by parts, and finally substitute back.



Two types of risks emerge:

- Skewness risk: risk that results when observations are not spread symmetrically around the average value. Ignoring this will cause the model to understate the risk of variables with high skewness.
- ▶ Kurtosis risk: kurtosis is associated with a high level of risk for an investment because it indicates that there are high probabilities of extremely large and extremely small returns.

Not taking them into account can thus severely impare our model.

Example. A gambler wants to know the expected value of the outcome of throwing two dice simultaneously.

- \triangleright He writes down all possible outcomes: $x_k \in \{2, 3, ..., 12\}$.
- \triangleright He determines the probabilities $f(x_k)$ as all possibilities to achieve x_k divided by the total number of outcomes, which is $6^2 = 36$. We get

$$f(2) = \frac{1}{36} = 0.02\overline{7}$$

$$f(7) = \frac{6}{36} = 0.1\overline{6}$$

$$f(3) = \frac{2}{36} = 0.0\overline{5}$$

$$f(8) = \frac{5}{36} = 0.13\overline{8}$$

$$f(4) = \frac{3}{36} = 0.08\overline{3}$$

$$\vdots$$

$$f(5) = \frac{4}{36} = 0.\overline{1}$$

$$f(11) = \frac{2}{36} = 0.0\overline{5}$$

$$f(6) = \frac{5}{36} = 0.13\overline{8}$$

$$f(12) = \frac{1}{36} = 0.02\overline{7}$$

 \triangleright Then he computes all values $x_k f(x_k)$. The first values are:

$$x_1 f(x_1) = 2 \cdot f(2) = 0.0\overline{5}$$

 $x_2 f(x_2) = 3 \cdot f(3) = 0.1\overline{6}$
 $x_3 f(x_3) = 4 \cdot f(4) = 0.\overline{3}$,

and so on.

▶ He sums up across all events and calculates the mean as

$$\mu = \sum_{k=1}^{12} x_k f(x_k) = 7.$$

He notes, that this is also the median, since the distribution is symmetrical.

Exercise. By now, he is so intrigued by all these calculations, that he does not want to stop there. He decides to also calculate the variance, the skewness, and the kurtosis. He wonders if you can check and verify his results:)

- All terms $(x_k \mu)^2 f(x_k)$ sum up to $\sum (x_k \mu)^2 f(x_k) = 5.833$ or, taking the square root, $\sigma = 2.4152$.
- All terms $(x_k \mu)^3 f(x_k)$ sum up to zero, since for each entry with a positive deviation $(x_k \mu)^3$ there is an identical one with a negative sign and with the same probability. Hence, the skewness is zero.
- All terms $(x_k \mu)^4 f(x_k)$ sum up to 80.5. dividing by $\sigma^4 = 34.0278$, he calculates the kurtosis of 2.3657.

We now extend the notion of a cumulative distribution function to multiple random variables.

For two random variables X and Y we define the *joint* (bivariate) distribution function as

$$F_{X,Y}(x,y) = \mathbb{P}[X \le x, Y \le y] = \mathbb{P}[\{X \le x\} \cap \{Y \le y\}]$$
$$= \mathbb{P}[\{\omega \in \Omega \mid X(\omega) \le x \text{ and } Y(\omega) \le y\}].$$

The joint probability density function $f_{X,Y}$ for two absolutely continuous random variables is defined as the derivative with respect to both arguments,

$$f_{X,Y}(x,y) = \frac{\partial^2 F_{X,Y}(x,y)}{\partial x \partial y}$$
.

Notice that

$$F_{X,Y}(x,y) = \int_{-\infty}^{x} \int_{-\infty}^{y} f_{X,Y}(s,t) \,dsdt$$

Naturally, we can define the joint cdf for any finite number of random variables X_1, \ldots, X_n ,

$$F_{X_1,\ldots,X_n}(x_1,\ldots,x_n) = \mathbb{P}[X_1 \leq x_1,\ldots,X_n \leq x_n].$$

The joint density is similarly defined as

$$f_{X_1,\ldots,X_n}(x_1,\ldots,x_n)=\frac{\partial^n F_{X_1,\ldots,X_n}(x_1,\ldots,x_n)}{\partial x_1\cdots\partial x_n},$$

and again we have

$$F_{X_1,\ldots,X_n}(x_1,\ldots,x_n)=\int_{-\infty}^{x_1}\cdots\int_{-\infty}^{x_n}f_{X_1,\ldots,X_n}(t_1,\ldots,t_n)\,\mathrm{d}t_1\cdots\mathrm{d}t_n\,.$$

We say two elements A and B of a σ -algebra are *independent* and we write $A \perp B$ if

$$\mathbb{P}[A \cap B] = \mathbb{P}[A]\mathbb{P}[B].$$

To see why we call this property independence, we recall the definition of the conditional probability and use 'independence',

$$\mathbb{P}[A \mid B] = \frac{\mathbb{P}[A \cap B]}{\mathbb{P}[B]} = \frac{\mathbb{P}[A]\mathbb{P}[B]}{\mathbb{P}[B]} = \mathbb{P}[A].$$

So the probability of A stays the same independent of the occurrence of B. The same is true the other way around

$$\mathbb{P}[B \mid A] = \frac{\mathbb{P}[B \cap A]}{\mathbb{P}[A]} = \mathbb{P}[B].$$

This notion of independence can now easily be translated to the independence of two random variables.

We say two random variables X and Y are independent if for all $x, y \in \mathbb{R}$ the sets $\{X \le x\}$ and $\{Y \le y\}$ are independent.

From here we directly get the equivalent property that the joint cdf is the product of the individual cdf's. That is, for all $x, y \in \mathbb{R}$ we have

$$F_{X,Y}(x,y) = F_X(x)F_Y(y).$$

This is also equivalent (if all densities exists) that for all $x, y \in \mathbb{R}$, we have

$$f_{X,Y}(x,y)=f_X(x)f_Y(y).$$

For the independence of *more than two random variables* we differentiate between *pairwise independence* and *mutual independence*.

We say a set $\{X_1, \ldots, X_n\}$ of random variables is *pairwise independent* if every pair of random variables is independent (as defined above).

However, we say they are *mutually independent* if for all $x_1, \ldots, x_n \in \mathbb{R}$ we have

$$F_{X_1,\ldots,X_n}(x_1,\ldots,x_n)=F_{X_1}(x_1)\cdot\ldots\cdot F_{X_n}(x_n).$$

s:fi

As already mentioned, if two random variables X and Y are independent, then they are also uncorrelated. A short calculation shows that,

$$\mathbb{E}[XY] = \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} xy \, f_{X,Y}(x,y) \, \mathrm{d}x \, \mathrm{d}y$$

$$= \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} xy \, f_X(x) f_Y(y) \, \mathrm{d}x \, \mathrm{d}y$$

$$= \int_{-\infty}^{\infty} x f_X(x) \left(\int_{-\infty}^{\infty} y f_Y(y) \, \mathrm{d}y \right) \, \mathrm{d}x$$

$$= \left(\int_{-\infty}^{\infty} x f_X(x) \, \mathrm{d}x \right) \left(\int_{-\infty}^{\infty} y f_Y(y) \, \mathrm{d}y \right)$$

$$= \mathbb{E}[X] \mathbb{E}[Y] .$$

Hence, $cov(X, Y) = \mathbb{E}[XY] - \mathbb{E}[X]\mathbb{E}[Y] = 0$ and hence $\rho_{X,Y} = 0$.

However, if two random variables X and Y are uncorrelated, then they do not have to be independent.

One of the standard example is the following. Consider $X \sim \mathcal{U}(-1,1)$ and define $Y = X^2$. Clearly, X and Y are *not independent!* But there correlation is 0, so there uncorrelated,

$$\begin{aligned} \operatorname{cov}(X,Y) &= \mathbb{E}[XY] - \mathbb{E}[X]\mathbb{E}[Y] = \mathbb{E}[X^3] - \mathbb{E}[X]\mathbb{E}[X^2] \\ &= \mathbb{E}[X^3] = \int_{-1}^{1} \frac{1}{2} x^3 dx = 0. \end{aligned}$$

Exercise. Come up with an own example of uncorrelated but dependent random variables.



Exercise. We want to practice working with distributions by means of the log-normal distribution. The log-normal distribution is often use to model returns of stocks.

We say X is log-normally distributed and write $X \sim \log \mathcal{N}(\mu, \sigma^2)$ if $Y = \log(X) \sim \mathcal{N}(\mu, \sigma^2)$

- \triangleright First of all, show that $X \sim \log \mathcal{N}(\mu, \sigma^2)$ can be written as $X = \exp(\mu + \sigma Z)$, where $Z \sim \mathcal{N}(0, 1)$.
- Now show that $\mathbb{E}[X] = \exp(\mu + \frac{\sigma^2}{2})$. You will have to calculate $\mathbb{E}[e^{\sigma Z}]$. For this write down the integral and *complete the square*. Try using $(z - \sigma)^2$.
- Derive the density function of X. Notice that it is only defined on $(0, +\infty)$. You will get

$$f(x) = \frac{1}{x\sqrt{2\pi\sigma^2}} \exp\left[-\frac{(\log x - \mu)^2}{2\sigma^2}\right]$$

Appendix

Maximum Likelihood (ML) methods





How are the parameters estimated from historical data in the models we have been considering?

A commonly applied approach is known as the *maximum likelihood (ML)* method.

It involves choosing values for the parameter that maximise the chance (or likelihood) of the data occurring.

Suppose we have a sample x_1, x_2, \dots, x_N of N i.i.d. random variables, coming from a parametric model.

▶ The joint density function of the observations is

$$f(x_1, x_2, \ldots, x_N | \theta) = f_1(x_1 | \theta) \cdot f_2(x_2 | \theta) \cdot \ldots \cdot f_n(x_N | \theta),$$

where θ summarises the model parameters.

- \triangleright The idea of the maximum likelihood (ML) method is to *choose* θ such that the joint density function is maximised, given the observed sample of data.
- A natural tool to this end is the *likelihood function*, which we define as

$$\mathcal{L}(\theta|x_1,x_2,\ldots,x_N):=f(x_1,x_2,\ldots,x_N|\theta)=\prod_{i=1}^N f_i(x_i|\theta).$$

To convert the product to summation (which is easier to handle on a computer), we take the logarithm. The result is called the *log-likelihood*:

$$\log \mathcal{L}(\theta|x_1,x_2,\ldots,x_n) = \sum_{i=1}^n \log(f_i(x_i|\theta)).$$

The ML method estimates θ by finding a value for θ that maximises $\log \mathcal{L}(\theta|x_1, x_2, \dots, x_n)$, that is,

$$\hat{\theta}_{\mathsf{mle}} := rg \max_{\theta} \log \mathcal{L}(\theta|x_1, x_2, \dots, x_n).$$

Estimate the variance of a variable X from m observations on X when the underlying distribution is *normal with zero mean*.

Let $u_1, u_2, ...$ denote the sample of m observations and denote the unknown variance parameter by v.

The likelihood of u_i being observed is the probability density function for X when $X = u_i$

$$\frac{1}{\sqrt{2\pi v}}\exp\left(\frac{-u_i^2}{2v}\right).$$

The likelihood of m observations occurring in order in which they are observed is

$$\prod_{i=1}^{m} \left[\frac{1}{\sqrt{2\pi \nu}} \exp\left(\frac{-u_i^2}{2\nu}\right) \right].$$

Using the maximum likelihood method, the best estimate of v is the value that maximises this expression.

Maximising an expression is equivalent to maximising the logarithm of that expression, since the logarithm is strictly increasing.

Taking logarithms and ignoring constant multiplicative factors, it can be seen that we wish to maximise

$$\sum_{i=1}^{m} \left[-\log v - \frac{u_i^2}{v} \right].$$

Differentiating this expression with respect to ν and setting the result equation to zero, we see that the maximum likelihood estimator of v is

$$\frac{1}{m}\sum_{i=1}^m u_i^2.$$

The natural starting point for learning about statistical data analysis is with a sample of independent identically distributed (i.i.d.) data, say $Y = (Y_1, \dots, Y_n)$, as for example in an idealized experiment of randomly drawing a marble from an urn.

The *linear regression model* relaxes both assumptions

- \triangleright allowing the means of the Y_i to depend, in a linear way, on other additional variables,
- \triangleright allowing for the Y_i to have different variances, and
- \triangleright allowing for correlation between the Y_i .

The linear regression model:

- is of fundamental importance in a large variety of quantitative disciplines, and
- it forms the basis of many complex and seemingly unrelated models.

The univariate linear regression model relates the scalar random variable Y to k other (possibly random) variables, or regressors, x_1, \ldots, x_k as

$$Y = \beta_1 x_1 + \beta_2 x_2 + \dots + \beta_k x_k + \epsilon$$
, where, typically, $\epsilon \sim \mathcal{N}\left(0, \sigma^2\right)$.

Values β_1, \ldots, β_k and σ^2 are unknown, constant parameters to be estimated from the data.

To emphasise that the means of the Y_i are not constant, we write

$$Y_i = \beta_1 x_{i,1} + \beta_2 x_{i,2} + \cdots + \beta_k x_{i,k} + \epsilon_i, \quad i = 1, 2, \dots, n.$$

The ϵ_i represent the difference between the values of Y_i and the model $\sum_{i=1}^{k} \beta_i x_{i,j}$, and so are referred to as the *error terms*.

Notice that the error terms are i.i.d., but the Y_i are not. However, if we take k = 1 and $x_{i,1} \equiv 1$, then the above equation reduces to

$$Y_i = \beta_1 + \epsilon_i \stackrel{\text{i.i.d.}}{\sim} \mathcal{N}\left(\beta_1, \sigma^2\right)$$
.

In fact, it is usually the case that $x_{i,1} \equiv 1$ for any $k \geq 1$, in which case the model is said to include a constant or have an intercept term.

In certain applications, the ordering of the dependent variable and the regressors is important, because they are observed in time, usually equally spaced. For this we use the notation Y_t , we get

$$Y_t = \beta_1 x_{t,1} + \beta_2 x_{t,2} + \dots + \beta_k x_{t,k} + \epsilon_t, \quad t = 1, 2, \dots, T.$$

An important special case is with k = 2, $x_{t,1} = 1$, and $x_{t,2}$ not constant.

Then

$$Y_t = \beta_1 + \beta_2 X_t + \epsilon_t, \quad t = 1, \dots, T,$$

is referred to as the simple linear regression model.

The goal is to find β_1 and β_2 which provide a 'best' fit. One possibility is to minimise the square-error, that is minimise the function

$$S(\beta_1, \beta_2) = \sum_{t=1}^{T} \epsilon_t^2 = \sum_{t=1}^{T} (Y_t - \beta_1 - \beta_2 X_t)^2.$$

For this, we set the partial derivatives equal zero, and solve for β_1 and β_2 ,

$$\partial_{\beta_1} S(\beta_1, \beta_2) = -2 \sum_{t=0}^{T} (Y_t - \beta_1 - \beta_2 X_t) \stackrel{!}{=} 0$$
, hence, $\hat{\beta}_1 = \bar{Y} - \beta_2 \bar{X}$.

Substituting $\hat{\beta}_1$ in S and differentiating with respect to β_2 gives us

$$\begin{aligned} \partial_{\beta_2} S(\hat{\beta}_1, \beta_2) &= \partial_{\beta_2} \left(\sum_{t=1}^{I} (Y_t - \bar{Y} - \beta_2 (X_t - \bar{X}))^2 \right) \\ &= -2 \sum_{t=1}^{T} (Y_t - \bar{Y} - \beta_2 (X_t - \bar{X})) (X_t - \bar{X}) \stackrel{!}{=} 0 \,, \end{aligned}$$

which leads us to

$$\hat{\beta}_2 = \frac{\sum_{t=1}^{T} (Y_t - \bar{Y})(X_t - \bar{X})}{\sum_{t=1}^{T} (X_t - \bar{X})^2} = \frac{\text{Cov}(X, Y)}{\text{Var}(X)}.$$

So we have

$$\hat{eta}_1 = ar{Y} - rac{\mathrm{Cov}(X,Y)}{\mathrm{Var}(X)}ar{X} \quad ext{and} \quad \hat{eta}_2 = rac{\mathrm{Cov}(X,Y)}{\mathrm{Var}(X)} \,.$$

In order to verify that this indeed provides a minimum, we consider the second partial derivatives,

$$\partial_{eta_1eta_1}S=2T>0\,,\;\;\partial_{eta_2eta_2}S=2\sum_{t=1}^IX_t^2>0\,,\;\; ext{and}$$
 $\partial_{eta_1eta_2}S=\partial_{eta_2eta_1}S=2\sum_{t=1}^TX_t\,.$

Hence, $(\partial_{\beta_1\beta_1}S)(\partial_{\beta_2\beta_2}S) - (\partial_{\beta_1\beta_2}S)^2 = 4(T\sum_{t=1}^T X_t^2 - (\sum_{t=1}^T X_t)^2) > 0$, confirming a minimum.

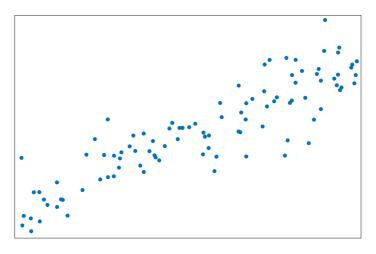


Figure 1: Linearly correlated data and the simple linear regression model

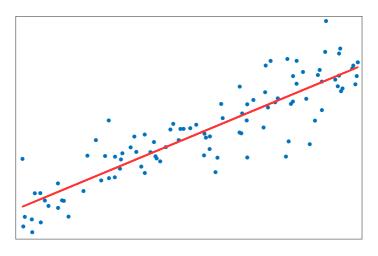


Figure 2: Linearly correlated data and the simple linear regression model

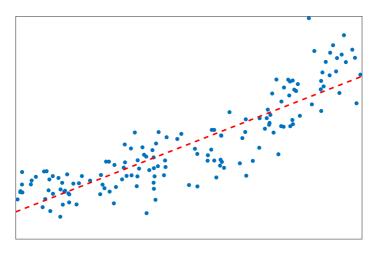


Figure 3: Nonlinearly correlated data and the simple linear regression model

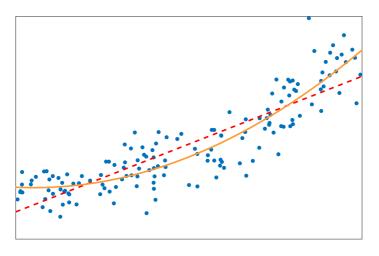


Figure 4: Nonlinearly correlated data and linear regression with three regressors

General linear regression model

We see that the simple linear regression model quickly reaches its limits. So we want to consider the general case.

Using standard matrix notation, with $Y = (Y_1, ..., Y_T)'$, $\beta = (\beta_1, ..., \beta_k)'$, $x_t = (x_{t,1}, ..., x_{t,k})'$,

$$\mathsf{X} = \begin{bmatrix} \mathsf{X}_1' \\ \vdots \\ \mathsf{X}_T' \end{bmatrix} = \begin{bmatrix} \mathsf{X}_{1,1} & \mathsf{X}_{1,2} & \cdots & \mathsf{X}_{1,k} \\ \mathsf{X}_{2,1} & \mathsf{X}_{2,2} & \cdots & \mathsf{X}_{2,k} \\ \vdots & \vdots & & \vdots \\ \mathsf{X}_{T,1} & \mathsf{X}_{T,2} & & \mathsf{X}_{T,k} \end{bmatrix}, \quad \text{and } \epsilon \sim \mathcal{N}\left(0, \sigma^2 \mathsf{I}_T\right),$$

we can compactly express the model as

$$Y = X\beta + \epsilon$$
.

Notice that $Y \sim \mathcal{N}\left(X\beta, \sigma^2I_T\right)$.

The method we used before is the most popular way of estimating the k parameters in β , namely, the *method of least squares*, which takes $\hat{\beta} = \arg\min S(\beta)$, where

$$S(\beta) = S(\beta; Y, X) = (Y - X\beta)'(Y - X\beta) = \sum_{t=1}^{T} (Y_t - x_t'\beta)^2.$$

This is referred to as *ordinary* least squares.

Assume that X is of full rank k. Using matrix calculus, we get

$$\partial S(\beta)/\partial \beta = -2X'(Y-X\beta),$$

and setting this to zero yields the solution

$$\widehat{\boldsymbol{\beta}} = \left(\mathsf{X}'\mathsf{X} \right)^{-1} \mathsf{X}'\mathsf{Y}.$$

This is referred to as the *ordinary least squares* (o.l.s.) estimator of β .

To verify that $\widehat{\beta} = (X'X)^{-1}X'Y$ indeed corresponds to the minimum of $S(\beta)$, the second derivative is checked for positive definiteness, yielding

$$\partial^2 S(\beta) / \partial \beta \partial \beta' = 2X'X,$$

which is necessarily positive definite when X is full rank.

Observe that, if X = 1, then

$$\widehat{eta}=(1'1)^{-1}1'\mathsf{Y}=rac{1}{T}\sum_{t=1}^T Y_t=ar{Y}, ext{ the mean of the } Y_t.$$

Also, if k = T (and X is full rank), then

$$\hat{\beta} = (X'X)^{-1} X'Y = X^{-1} (X')^{-1} X'Y = X^{-1}Y$$

and using this we get
$$S(\widehat{\beta}) = (Y - X\widehat{\beta})'(Y - X\widehat{\beta}) = 0.$$

Notice that the derivation of $\widehat{\beta}$ did not involve any explicit distributional assumptions. Consequently, the estimator may not have any meaning if the maximally existing moment of the $\{\epsilon_t\}$ is too low.

- ightharpoonup Take for example X = 1 and let $\epsilon_t \stackrel{\mathrm{i.i.d.}}{\sim} \mathrm{Cauchy}$. Then $\widehat{\beta} = \overline{Y}$ is a useless estimator.
- ightharpoonup Assume that the first moment of the $\{\epsilon_t\}$ exists and is zero. Writing

$$\widehat{\boldsymbol{\beta}} = \left(\mathsf{X}'\mathsf{X} \right)^{-1} \mathsf{X}' \left(\mathsf{X} \boldsymbol{\beta} + \boldsymbol{\epsilon} \right) = \boldsymbol{\beta} + \left(\mathsf{X}'\mathsf{X} \right)^{-1} \mathsf{X}' \boldsymbol{\epsilon},$$

we see that $\widehat{oldsymbol{eta}}$ is *unbiased*, which means that

$$\mathbb{E}[\widehat{\boldsymbol{\beta}}] = \boldsymbol{\beta} + \left(\mathsf{X}'\mathsf{X}\right)^{-1}\mathsf{X}'\mathbb{E}[\boldsymbol{\epsilon}] = \boldsymbol{\beta}.$$

Next, if we have existence of second moments, and $Var(\epsilon) = \sigma^2 I$, then

$$\begin{aligned} \operatorname{Var}(\widehat{\boldsymbol{\beta}} \mid \sigma^2) &= \mathbb{E}[(\widehat{\boldsymbol{\beta}} - \boldsymbol{\beta})(\widehat{\boldsymbol{\beta}} - \boldsymbol{\beta})' \mid \sigma^2] \\ &= \left(\mathsf{X}'\mathsf{X} \right)^{-1} \mathsf{X}' \mathbb{E}\left[\boldsymbol{\epsilon} \boldsymbol{\epsilon}' \right] \mathsf{X} \left(\mathsf{X}'\mathsf{X} \right)^{-1} = \sigma^2 \left(\mathsf{X}'\mathsf{X} \right)^{-1}. \end{aligned}$$

The *Gauss-Markov Theorem* tells us that $\widehat{\beta}$ has the *smallest variance* among all linear unbiased estimators. We call such an estimator BLUE – best linear unbiased estimator.

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The *Gauss-Markov Theorem* tells us that $\widehat{\beta}$ has the *smallest variance* among all linear unbiased estimators. We call such an estimator BLUE – best linear unbiased estimator.

To see this, consider another linear estimator $\widehat{\boldsymbol{\beta}}^* = A'Y$ and let $D = A - X(X'X)^{-1}$. The unbiased property $\mathbb{E}\big[\widehat{\boldsymbol{\beta}}^*\big] \stackrel{!}{=} \boldsymbol{\beta}$ implies that D'X = 0.

Next, we calculate $\mathrm{Var}(\widehat{\boldsymbol{\beta}}^* \mid \sigma^2) = \mathrm{Var}(\widehat{\boldsymbol{\beta}} \mid \sigma^2) + \sigma^2 D' D$. The result follows because D'D is positive semi-definite and the variance is minimized when D = 0.

In many situations, it is reasonable to assume normality for the $\{\epsilon_t\}$. In this case, we can estimate σ^2 and β_i , $i=1,\ldots,k$, by maximum likelihood.

We start with the density function

$$f_{Y}(y) = (2\pi\sigma^{2})^{-T/2} \exp\Big\{-\frac{1}{2\sigma^{2}}\underbrace{(y-X\beta)'(y-X\beta)}_{=S(\beta;y,X)}\Big\},$$

and the log-likelihood

$$\ell\left(\boldsymbol{\beta}, \sigma^2; \mathsf{Y}\right) = -\frac{T}{2}\log\left(2\pi\right) - \frac{T}{2}\log\left(\sigma^2\right) - \frac{1}{2\sigma^2}\mathsf{S}(\boldsymbol{\beta}).$$

To get the Maximum Likelihood Estimator (m.l.e.) we have to solve

$$\frac{\partial \ell}{\partial \beta} = -\frac{2}{2\sigma^2} \mathsf{X}' \left(\mathsf{Y} - \mathsf{X}\beta \right) \stackrel{!}{=} 0$$

and

$$\frac{\partial \ell}{\partial \sigma^2} = -\frac{T}{2\sigma^2} + \frac{1}{2\sigma^4} \,\mathsf{S}(\beta) \stackrel{!}{=} 0.$$

This yields the same estimator for β and the m.l.e. of σ^2 , namely

$$\widehat{oldsymbol{eta}} = \left(\mathsf{X}'\mathsf{X} \right)^{-1} \mathsf{X}'\mathsf{Y} \quad \mathsf{and} \quad \widetilde{\sigma}^2 = \left. \mathsf{S}(\widehat{oldsymbol{eta}}) \middle/ \mathcal{T}. \right.$$

Let X, Y real valued random variables with finite second moment. Then:

$$\mathbb{E}[XY] \leq \mathbb{E}[X]^2 \mathbb{E}[Y]^2,$$

with equality if $X = \alpha Y$ for some $\alpha \in \mathbb{R}$.

Let x be a column vector. Then the following rules apply:

$$\frac{a'x}{\partial x} = \frac{\partial x'a}{\partial x} = a',$$

$$\frac{\partial b'Ax}{\partial x} = b'A,$$

$$\frac{\partial x'Ax}{\partial x} = x'(A + A'),$$

$$\frac{\partial g(u)}{\partial x} = \frac{\partial g(u)}{\partial u} \frac{\partial u}{\partial x},$$

where a, b are vectors that are not functions of x, A is a squared matrix that is not a function of x and u = u(x).

References

▶ Marc S. Paolella, *Linear Models and Time-Series Analysis*: Regression, ANOVA, ARMA and GARCH

Acknowledgements





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